THE FAR-FIELD MODELLING OF TRANSONIC COMPRESSIBLE FLOWS

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Dedicated to Professor Dr. J. Nečas on the occasion of his 70th birthday

Abstract. We present a method for the construction of artificial far-field boundary conditions for two- and three-dimensional exterior compressible viscous flows in aerodynamics. Since at some distance to the surrounded body (e.g. aeroplane, wing section, etc.) the convective forces are strongly dominant over the viscous ones, the viscosity effects are neglected there and the flow is assumed to be inviscid. Accordingly, we consider two different model zones leading to a decomposition of the original flow field into a bounded computational domain (near field) and a complementary outer region (far field). The governing equations as e.g. compressible Navier-Stokes equations are used in the near field, whereas the inviscid far field is modelled by Euler equations linearized about the free-stream flow. By treating the linear model analytically and numerically, we get artificial far-field boundary conditions for the (nonlinear) interior problem. In the two-dimensional case, the linearized Euler model can be handled by using complex analysis. Here, we present a heterogeneous coupling of the above-mentioned models and show some results for the flow around the NACA0012 airfoil. Potential theory is used for the three-dimensional case, leading also to non-local artificial far-field boundary conditions.

Keywords: artificial boundary and transmission conditions, compressible transonic flow, linearized Euler equations, integral equations with kernels of Cauchy type, potential theory, domain decomposition

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1. Introduction

A sensitive issue in the numerical treatment of exterior flow problems is the accurate description of the flow field also far away from the travelling bodies. In various problems in CFD, it turns out that the influence of the far field on the flow
properties in the immediate vicinity of the body is significant. Constant or characteristic boundary conditions involving constant values at infinity are simple and widely used in transonic flow computations. However, such conditions need to be imposed far enough from the object (a requirement which may lead to tremendous computational costs, even when using unstructured meshes), and can also generate nondesirable reflections at the far-field boundary, see e.g. Hirsch [10].

A possible incorporation of the far-field behaviour, by maintaining the efficiency of the numerical discretization, is to replace in the far-field regions the (mostly expensive) complete model and the (rather inaccurate) constant flow by intermediate models, obtained e.g. by neglecting terms which are not relevant there and by linearization.

The far-field simplifications should be chosen in such a way that on the one hand the outer flow is still modelled accurately enough, while on the other hand, the computational times are reasonable. The coupling of different models leads to heterogeneous domain decomposition methods, efficiently used in the numerical treatment of exterior boundary value problems, see e.g. [8], [15], [11], [2], [1] for problems in aerodynamics.

The physical properties of compressible Navier-Stokes flows around 2D or 3D bodies, and a dimensional analysis of different quantities in the equations indicate that the viscous terms are of order of magnitude comparable with the convective terms only inside a thin region near to the body, whereas away from it, the shear stresses are extremely small and are strongly dominated by the convective part. Consequently, we neglect the viscous terms in some distance to the body and model the outer far-field domain with the Euler equations, linearized about the free-stream subsonic uniform background flow; this model can treat flows with wakes admitting non-zero vorticity and non-constant entropy at the outflow.

By solving the simplified model in the far field, we generate artificial far-field boundary conditions for the interior problem. The linearized Euler system is partly solved by using potential theory. The corresponding boundary conditions are based on integral representations of the solution, hence they are non-local. We present the construction of these conditions, which are exact with respect to the linearized model considered, and show results for the two-dimensional flow around the NACA0012 airfoil.
We introduce here two models for the description of the near-field region around some solid body, and of the corresponding far-field domain.

2. Compressible flow models

2.1. Mathematical modelling of the near field

Let \( \Omega_a \subset \mathbb{R}^d \) (\( d = 2,3 \)) be a bounded domain representing a given body (as e.g. aeroplane, wing section, etc.). The conservation laws lead to the continuity equation, the Navier-Stokes equations and the energy equation written as the system in conservative form:

\[
\frac{\partial w}{\partial t} + \sum_{i=1}^{d} \frac{\partial f_i(w)}{\partial x_i} = \sum_{i=1}^{d} \frac{\partial R_i(w, \nabla_x w)}{\partial x_i} \quad \text{in} \quad \Omega_a \times (0,T).
\]

Here \((0,T)\) is a time interval; \( w = (\rho, \rho v, e)^\top \) collects the conservative variables with \( \rho \) the density, \( v = (v_1, \ldots, v_d)^\top \) the velocity vector, and \( e \) the total energy. The inviscid fluxes \( f_i \) and the viscous terms \( R_i \) are given by

\[
\begin{align*}
 f_i(w) &= (\rho v_i, \rho v_i v_1 + \delta_{i1} p, \ldots, \rho v_i v_d + \delta_{id} p, (e + p)v_i)^\top, \\
 R_i(w, \nabla_x w) &= (0, \tau_{i1}, \ldots, \tau_{id}, \tau_{i1} v_1 + \ldots + \tau_{id} v_d + k \partial \theta / \partial x_i)^\top
\end{align*}
\]

for \( i = 1, \ldots, d \). The functions \( p, \theta \) and \( \tau_{ij} \) denote the pressure, the absolute temperature and the components of the viscous part of the stress tensor \( \tau_{ij} = \lambda \text{div} v \delta_{ij} + \mu (\partial v_i / \partial x_j + \partial v_j / \partial x_i) \) for \( i, j = 1, \ldots, d \), respectively. Here, \( \lambda \) and \( \mu \) are the viscosity coefficients and \( k \) is the heat conductivity. In addition, we employ the equations of state

\[
\begin{align*}
 p &= R \rho \theta \quad \text{and} \quad e = \rho \left( c_V \theta + \frac{|v|^2}{2} \right),
\end{align*}
\]

with the universal gas constant \( R \) and the specific heat \( c_V \) at constant volume, related by \( R = (\kappa - 1)c_V \), where \( \kappa \) is Poisson's adiabatic constant.

Existence and uniqueness results for compressible viscous flows with the Navier-Stokes equations exist only for small data and under restrictive assumptions (as e.g. ideal gas in thermal equilibrium). Even for simpler models, as e.g. the full potential equation, existence of transonic solutions is still open, see Feistauer and Nečas [7]. However, we are interested here in the numerical treatment of compressible viscous flows.

Since the viscous terms are of order of magnitude comparable with the convective terms only inside Prandtl’s boundary layer, and away from it, the viscous stress
tensor becomes strongly dominated by the inertial forces, we neglect these terms in some distance to the body and decompose the exterior computational domain $\Omega_1$ into a bounded domain $\Omega$, where we employ the full governing system of equations, and a corresponding far field $\Omega_2 := \mathbb{R}^d \setminus (\overline{\Omega} \cup \overline{\Omega}_o)$, where we assume the flow to be inviscid.

2.2. Mathematical modelling of the far field

We begin our analysis in the exterior $\Omega_2$ with the 3D Euler equations

\[
\begin{align*}
\frac{\partial (\rho v_1)}{\partial x_1} + \frac{\partial (\rho v_2)}{\partial x_2} + \frac{\partial (\rho v_3)}{\partial x_3} &= 0, \\
\frac{\partial v_1}{\partial x_1} + \frac{\partial v_2}{\partial x_2} + \frac{\partial v_3}{\partial x_3} + \frac{\partial p}{\partial x_1} &= 0, \\
\frac{\partial v_1}{\partial x_1} + \frac{\partial v_2}{\partial x_2} + \frac{\partial v_3}{\partial x_3} + \frac{\partial p}{\partial x_2} &= 0, \\
\frac{\partial v_1}{\partial x_1} + \frac{\partial v_2}{\partial x_2} + \frac{\partial v_3}{\partial x_3} + \frac{\partial p}{\partial x_3} &= 0, \\
v_1 \frac{\partial p}{\partial x_1} + v_2 \frac{\partial p}{\partial x_2} + \kappa p \left( \frac{\partial v_1}{\partial x_1} + \frac{\partial v_2}{\partial x_2} + \frac{\partial v_3}{\partial x_3} \right) &= 0
\end{align*}
\]

for the primitive variables $\rho, v_1, v_2, v_3$ and $p$. In $\Omega_2$ we represent these as

\[
(1) \quad \varrho = \varrho_\infty + \tilde{\varrho}, \quad v_1 = v_{1\infty}^{(1)} + \tilde{v}_1, \quad v_2 = v_{2\infty}^{(2)} + \tilde{v}_2, \quad v_3 = v_{3\infty}^{(3)} + \tilde{v}_3, \quad p = p_\infty + \tilde{p},
\]

where $V_\infty := (\varrho_\infty, v_{1\infty}^{(1)}, v_{2\infty}^{(2)}, v_{3\infty}^{(3)}, p_\infty)^\top$ defines the free-stream subsonic flow. To simplify our analysis, we specify the coordinate system such that

\[
(2) \quad \varrho_\infty \equiv (\varrho_\infty^{(1)}, 0, 0)^\top
\]

having in mind that the general case of non-zero $v_{1\infty}^{(2)}$ and $v_{3\infty}^{(3)}$ can be reduced to (2) by an elementary transformation. For the perturbation variables $V := (\tilde{\varrho}, \tilde{v}_1, \tilde{v}_2, \tilde{v}_3, \tilde{p})^\top$ we obtain via linearization of (5):

\[
(3) \quad \varrho_\infty v_{1\infty}^{(1)} \frac{\partial \tilde{\varrho}}{\partial x_1} + \varrho_\infty \left( \frac{\partial \tilde{v}_1}{\partial x_1} + \frac{\partial \tilde{v}_2}{\partial x_2} + \frac{\partial \tilde{v}_3}{\partial x_3} \right) = 0,
\]

\[
(4) \quad \varrho_\infty v_{2\infty}^{(1)} \frac{\partial \tilde{\varrho}}{\partial x_1} = 0, \quad \varrho_\infty v_{1\infty}^{(1)} \frac{\partial \tilde{v}_2}{\partial x_1} + \frac{\partial \tilde{p}}{\partial x_2} = 0, \quad \varrho_\infty v_{1\infty}^{(1)} \frac{\partial \tilde{v}_3}{\partial x_1} + \frac{\partial \tilde{p}}{\partial x_3} = 0,
\]

\[
(5) \quad v_{1\infty}^{(1)} \frac{\partial \tilde{p}}{\partial x_1} + \kappa p_\infty \left( \frac{\partial \tilde{v}_1}{\partial x_1} + \frac{\partial \tilde{v}_2}{\partial x_2} + \frac{\partial \tilde{v}_3}{\partial x_3} \right) = 0.
\]

Here we neglect terms of the form $\tilde{h}_1 \cdot \partial \tilde{h}_2 / \partial x_j, \partial (\tilde{h}_1 \tilde{h}_2) / \partial x_j (j = 1, \ldots, d)$, which are small of quadratic order, where $\tilde{h}_1, \tilde{h}_2 \in \{\tilde{\varrho}, \tilde{v}_1, \tilde{v}_2, \tilde{p}\}$ represent perturbations of the
flow variables. Hence, the linear model can describe inviscid flows characterized by small vorticity. Nevertheless, the linearized Euler equations are capable to model a “weakly” rotational, compressible flow, and their numerical treatment requires minimal computational cost. For subsonic free-stream flow, (8) is a system of composed type, having three real and two complex conjugate eigenvalues.

Let us denote by \( a_\infty := \sqrt{\frac{\kappa p}{\rho}} / a_\infty \) and \( M_\infty := |v_\infty| / a_\infty \) the local speed of sound and the Mach number at infinity, respectively. Now we introduce a new Cartesian system \((\xi_1, \xi_2, \xi_3)\) by

\[
(9) \quad \xi_1 = \hat{x}_1 / \sqrt{1 - M_\infty^2}, \quad \xi_2 = \hat{x}_2, \quad \xi_3 = \hat{x}_3 \quad \text{(Prandtl transformation)},
\]

and, in addition, new unknown functions (perturbations) by

\[
(10) \quad \bar{\rho} = \tilde{\rho} \cdot \frac{\kappa p}{\rho_\infty \sqrt{1 - M_\infty^2}} v_\infty^{(1)}, \quad \bar{v}_1 = \tilde{v}_1 \sqrt{1 - M_\infty^2}, \quad \bar{v}_2 = \tilde{v}_2, \quad \bar{v}_3 = \tilde{v}_3, \quad \bar{p} = \tilde{p} \cdot \frac{\sqrt{1 - M_\infty^2}}{\rho_\infty v_\infty^{(1)}}.
\]

With these transformations, the system can be reformulated to an elliptic subsystem for \( \bar{p}, \bar{v}_2 \) and \( \bar{v}_3 \), and two transport equations for \( \bar{\rho} - \bar{p} \) and \( \bar{v}_1 + \bar{p} \):

\[
\left\{ \begin{array}{l}
\frac{\partial \bar{v}_2}{\partial \xi_1} + \frac{\partial \bar{p}}{\partial \xi_2} (\bar{x}) = 0, \\
\frac{\partial \bar{v}_3}{\partial \xi_1} + \frac{\partial \bar{p}}{\partial \xi_3} (\bar{x}) = 0, \\
\frac{\partial}{\partial \xi_1} \left[ - \frac{\partial \bar{p}}{\partial \xi_1} + \frac{\partial \bar{v}_2}{\partial \xi_2} + \frac{\partial \bar{v}_3}{\partial \xi_3} \right] (\bar{x}) = 0, \\
\frac{\partial}{\partial \xi_1} (\bar{\rho} - \bar{p})(\bar{x}) = 0, \\
\frac{\partial}{\partial \xi_1} (\bar{v}_1 + \bar{p})(\bar{x}) = 0, \quad \text{for} \quad \bar{x} \in \Omega_2.
\end{array} \right.
\]

We require the flow quantities to be bounded at infinity and to tend in the upstream region to the free-stream quantities. Therefore, we supplement the system (11) with the condition

\[
(12) \quad |\bar{\rho}|, |\bar{v}_1|, |\bar{v}_2|, |\bar{v}_3|, |\bar{p}| < \infty \quad \text{for} \quad |\bar{x}| \rightarrow \infty \quad \text{and} \quad \bar{\rho}, \bar{v}_1, \bar{v}_2, \bar{v}_3, \bar{p} \rightarrow 0 \quad \text{as} \quad \bar{x}_1 \rightarrow -\infty.
\]

3. Far-field boundary conditions for 2D flows

In the two-dimensional case, the artificial far-field boundary separating the model zones \( \Omega_1 \) and \( \Omega_2 \) is a simply closed curve \( \Gamma_\infty \), which we assume to be piecewise smooth. In the sequel, \( \mathbf{n} \) denotes the unit normal vector on \( \partial \Omega_1 \) exterior to the interior domain \( \Omega_1 \).
For a given velocity field $\mathbf{v}$, the input and the output parts of the boundary $\Gamma_\infty$ are defined with respect to the far-field domain $\Omega_2$ by

$$
\Gamma_{\infty,\text{in}} := \{ \mathbf{x} \in \Gamma_\infty \mid (\mathbf{v} \cdot \mathbf{n})(\mathbf{x}) \geq 0 \},
$$

$$
\Gamma_{\infty,\text{out}} := \Gamma_\infty \setminus \Gamma_{\infty,\text{in}}.
$$

In two space dimensions, the system (11) of the linearized Euler equations does not contain equation (11)$_2$, and equation (11)$_3$ takes the form

$$(13) \quad \left[ -\frac{\partial \sigma}{\partial x_1} + \frac{\partial \sigma_2}{\partial x_2} \right](\mathbf{x}) = 0 \quad \text{for} \quad \mathbf{x} \in \Omega_2.$$

The elliptic subsystem, consisting here of equations (11)$_1$ and (13), is the Cauchy-Riemann system for $\sigma$ and $\sigma_2$ in $\Omega_2$ and can be solved by using the complex Cauchy integral. Let $\psi(\zeta) = \sigma^0(\zeta) + i\sigma_2^0(\zeta)$, $\zeta = \xi_1 + i\xi_2 \in \Gamma_\infty$ be a given complex-valued function defined on $\Gamma_\infty$. Then the function $\chi: \Omega_2 \rightarrow \mathbb{C}$, defined by the Cauchy potential

$$(14) \quad \chi(z) = \sigma(z) + i\sigma_2(z) := \frac{1}{2\pi i} \oint_{\Gamma_\infty} \frac{\psi(\zeta)}{\zeta - z} \, d\zeta \quad \text{for} \quad z = \xi_1 + i\xi_2 \in \Omega_2,$$

is a holomorphic solution of the elliptic subsystem (11)$_1$, (13). Since it is necessary to solve the linearized Euler model up to the boundary $\Gamma_\infty$, we need the behavior of the Cauchy potential $\chi(z)$ as $z \in \Omega_2$ approaches a given point $z_0$ situated on $\Gamma_\infty$.

**Lemma 3.1.** Let $\psi: \Gamma_\infty \rightarrow \mathbb{C}$, $\psi \in C^1(\Gamma_\infty)$ be the trace of some solution to the Cauchy-Riemann system in the exterior domain $\Omega_2$. Then, for arbitrary $z_0 \in \Gamma_\infty$, the value $\psi(z_0)$ satisfies the relation

$$(15) \quad \psi(z_0) = \frac{1}{\pi i} \lim_{\varepsilon \rightarrow 0} \oint_{\Gamma_\infty \cap \{ |\zeta - z_0| > \varepsilon \}} \frac{\psi(\zeta)}{\zeta - z_0} \, d\zeta,$$

and the Cauchy principal value exists.
Lemma 3.2 (Plemelj-Sokhotski). Let $\psi: \Gamma_\infty \to \mathbb{C}$ be Hölder-continuous on $\Gamma_\infty$. Then, for every $z_0 \in \Gamma_\infty$, the jump relation
\begin{align}
\lim_{z \to z_0} \chi(z) = \lim_{z \to z_0} \frac{1}{2\pi i} \oint_{\Gamma_\infty} \frac{\psi(\zeta)}{\zeta - z} \, d\zeta = \frac{1}{2\pi i} \lim_{\epsilon \to 0} \oint_{\Gamma_\infty \setminus |\zeta - z_0| \geq \epsilon > 0} \frac{\psi(\zeta)}{\zeta - z_0} \, d\zeta + \frac{1}{2} \psi(z_0)
\end{align}
holds.

Corollary. If $\psi(\zeta)$ is the trace of some solution to the Cauchy-Riemann equations in $\Omega_2$, then
\begin{align}
\lim_{z \to z_0} \chi(z) = \psi(z_0) \quad \text{for all } z_0 \in \Gamma_\infty.
\end{align}
Hence, the trace of the Hilbert transform (14) on $\Gamma_\infty$ is a projection.

This projection property will decisively be exploited for the numerical realization of the coupling [2]. It is valid also in the three-dimensional case; its interpretation will be given in detail in the next section.

For numerical treatment, we impose as transmission condition for the elliptic components $\rho$ and $\varphi$ their continuity across $\Gamma_\infty$:
\begin{align}
\varphi_{LE} :\rho_{NS} \quad \text{and} \quad \rho_{LE} :\varphi_{NS},
\end{align}
compute these quantities in the far field by using the Cauchy integral (14) (in the case of an overlapping scheme) or the Plemelj-Sokhotski formula to get their traces on $\Gamma_\infty$ (in the case of a non-overlapping scheme).

As transmission conditions for the equations (11)$_4$, we require the continuity of those characteristic variables which correspond to characteristic directions “entering” the domain of hyperbolicity (see e.g. Hirsch [10], Quarteroni and Stolc [11]). On the inflow boundary $\Gamma_\infty$, we impose the continuity of the Riemann invariants: The Riemann invariants corresponding to the linearized Euler equations in $\Omega_2$ are equated to their counterparts from the interior domain, i.e. to the Riemann invariants corresponding to the (nonlinear) 2D steady Euler equations in $\Omega_1$ (for more details, see Coclici and Wendland [3]). On the output boundary $\Gamma_\infty$, the characteristics are outgoing and therefore we impose here the compatibility conditions
\begin{align}
\frac{\partial}{\partial \varphi_2} (\varphi - \rho)(\varphi_1, \varphi_2) = 0 \quad \text{and} \quad \frac{\partial}{\partial \varphi_1} (\varphi_1 + \rho)(\varphi_1, \varphi_2) = 0.
\end{align}
We find the solution (11)$_4$ in the form
\begin{align}
\varphi(\varphi_1, \varphi_2) = \rho(\varphi_1, \varphi_2) + c_1(\varphi_2) \quad \text{and} \quad \varphi_1(\varphi_1, \varphi_2) = -\rho(\varphi_1, \varphi_2) + c_2(\varphi_2),
\end{align}
where \( p = p^{LE} \) is already known and where, as a consequence of (12), the functions \( c_1, c_2 \) are equal to zero in the upstream domain.

As interface conditions from \( \Omega_2 \) to \( \Omega_1 \) we impose the continuity of the normal fluxes across \( \Gamma_{\infty} \). The normal inviscid flux of the linearized Euler equations in \( \Omega_2 \) is set equal to the normal total flux for the Navier-Stokes system solution (see, e.g. Quarteroni and Stolcis [11]):

\[
(20) \quad -\sum_{i=1}^{2} R_i(w, \nabla w)n_i + \sum_{i=1}^{2} f_i(w)n_i = \sum_{i=1}^{2} f^{LE}_i(w)n_i \quad \text{across} \quad \Gamma_{\infty}.
\]

Here, the flux functions \( f^{LE}_i \) are linearizations of the Euler fluxes \( f_i \) about the free-stream flow parameters \( \mathbf{V}_{\infty} \).

For the numerical solution of the Navier-Stokes equations we use the operator-splitting FVM-FEM algorithm by M. Feistauer (Prague, see Feistauer et al. [6]) executed on a triangular mesh (Figure 1, left). We present results for flows around the NACA0012 airfoil at the free-stream Mach number \( M_{\infty} = 0.85 \), for \( Re = 5000 \) and zero angle of attack. In Figures 1 (right) and 2 we show the isolines of the Mach number, of the entropy and of the vorticity field corresponding to the coupled solution.

![Fig. 1 Computational domain and isolines of the Mach number](image)

4. Solution of the 3D linearized Euler system

In the three-dimensional case, the artificial far-field boundary \( S_{\infty} := \partial \Omega_1 \cap \partial \Omega_2 \) is a closed surface, which we assume to consist of a finite number of simply connected regular surface segments \( S_{\infty,i} \), i.e. \( S_{\infty} = \bigcup_i S_{\infty,i} \) with \( S_{\infty,i} \in C^1 \) for \( i = 1, \ldots, m \).
We begin our analysis with the elliptic subsystem (11)$_{1,2,3}$ and collect the elliptic variables into the vector-valued function

\[ \mathbf{E}(\vec{x}) := (-\vec{v}_1, \vec{v}_2, \vec{v}_3)^\top(\vec{x}) \text{ for } \vec{x} \in \Omega_2, \]

and add to (11)$_{1,2,3}$ the equation

\[ \frac{\partial \vec{v}_3}{\partial x_2} - \frac{\partial \vec{v}_2}{\partial x_3} = b \]

with an auxiliary function \( b: \Omega_2 \rightarrow \mathbb{R} \). Then (11)$_{1,2,3}$, (22) take the form

\[ \begin{cases} \text{rot} \mathbf{E} = b, \\ \text{div} \mathbf{E} = 0 \text{ in } \Omega_2 \end{cases} \]

with \( b := (b, 0, 0)^\top: \Omega_2 \rightarrow \mathbb{R}^3 \) to be specified later on. From (12) we get the conditions at infinity

\[ \lim_{|\vec{x}| \to \infty} |\mathbf{E}(\vec{x})| < \infty, \quad \lim_{x_1 \to \infty} \mathbf{E}(\vec{x}) = 0. \]

In what follows, we are concerned with the solution of the exterior boundary value problem (23), (24). Since we have to model also the wake, let us introduce the cylindrical “shadow” domain \( \Omega_W \subset \Omega_2 \) by

\[ \Omega_W := \{ \vec{x} = (x_1, x_2, x_3)^\top \in \Omega_2; \ x_1 > \bar{y}_1, x_2 = \bar{y}_2, x_3 = \bar{y}_3, \bar{y} = (\vec{y}_1, \vec{y}_2, \vec{y}_3)^\top \in \Omega_1 \} \]

and define the “outflow” part \( \omega = S_\infty \cap \Omega_W \) of the artificial boundary \( S_\infty \), see Figure 3.
Assuming that $\omega$ has a single-valued projection on the plane $\mathbb{R}_1 = 0$, we obtain

**Lemma 4.1.** The function $b$, associated with the solution of the linearized Euler equations via (22), satisfies the relations

\[
\begin{align*}
    b &\equiv 0 \text{ in } \Omega_2 \setminus \Omega_W, \\
    b &\equiv b(\mathbb{R}_2, \mathbb{R}_3) \text{ in } \Omega_W, \\
    \int_\omega b \cdot d\sigma &= 0.
\end{align*}
\]

Consequently, the specific properties of the right-hand side $b$ in (23) restrict the set of admissible vector fields $E$, which permits us to model the far-field behaviour. In order to proceed with our analysis, we introduce for arbitrary vector fields $f \in C(S_\infty)$ with $f|_{S_\infty,i} \in C^1(S_\infty,i)$ the surface-curl quantity

\[
(n \times \nabla, f)(\bar{x}) := \left[ \sum_{k=1}^3 (n \times \nabla) f_k \right](\bar{x})
\]

containing only the tangential derivatives of $f$.

Let us consider an arbitrary vector field $a: S_\infty \to \mathbb{R}^3$ with $a \in C(S_\infty)$ and $a|_{S_\infty,i} \in C^1(S_\infty,i)$ ($i = 1, \ldots, m$), satisfying with $b_n = b_n(a) := (n \times \nabla, a): S_\infty \to \mathbb{R}$ the conditions

\[
\text{supp}(b_n) \subset \omega \quad \text{and} \quad \int_\omega b_n d\sigma = 0.
\]

In the sequel we shall assume the general case $b_n \neq 0$ which corresponds to a non-zero wake. Denote by $e_1, e_2, e_3$ the basis vectors of the cartesian coordinate system. In dependence on $a$, we define now the vector field $b: \Omega_2 \to \mathbb{R}^3$ by

\[
b(\bar{x}) := \begin{cases} 
    \left( \frac{b_n(\mathbb{R}_1, \mathbb{R}_2, \mathbb{R}_3)}{(n, e_1)}, 0, 0 \right) & \text{for } \bar{x} \in \Omega_W, \\
    0 & \text{for } \bar{x} \in \Omega_2 \setminus \Omega_W.
\end{cases}
\]
In what follows, we rename for brevity the variables $\bar{x} \to x$. By means of the vector-valued functions $a: S_\infty \to \mathbb{R}^3$ and $b = b(a): \Omega_2 \to \mathbb{R}^3$, we introduce the vector field $E = E(x): \Omega_2 \to \mathbb{R}^3$ by

$$E(x) := -\int_\omega \frac{1}{2} \left( 1 + \frac{x_1 - y_1}{|x - y|} \right) (n \times \nabla_x a) e_1 \times \nabla_y \varphi(x' - y') d\sigma_y$$

$$-\int_{S_\infty} \left\{ [n \times a] \times \nabla_y \Phi(x - y) + (n \cdot a) \nabla_y \Phi(x - y) \right\} dS_y,$$

(28)

where $\Phi(x - y) := -\frac{1}{4\pi |x - y|}$ and $\varphi(x' - y') := -\frac{1}{2\pi} \log |x' - y'|$ denote the fundamental solutions of the Laplacian in three and two dimensions, respectively, and where $x' := (x_2, x_3) \in \mathbb{R}^2$ for $x \in \mathbb{R}^3$. The solution of the linearized Euler system in the far-field region $\Omega_2$ can now be characterized in the following theorems (see Sofronov and Wendland [14]).

**Theorem 4.2.** For arbitrary $a$ satisfying (26), the function $E$ evaluated via (28) is a solution of the system

$$\begin{cases}
\text{rot } E = b, \\
\text{div } E = 0
\end{cases}$$

with $b$ defined by (27). Moreover, the following estimate is valid for $E \equiv (-\varphi, \tau_2, \tau_3)^T$ as $|x| \to \infty$ (here $\vartheta$ denotes the angle between $x$ and $e_1$, see Figure 3):

$$|\varphi| = O(|x|^{-2}) \quad \text{and} \quad |\tau_2|, |\tau_3| = O \left( \frac{1}{1 + |x|^2 \sin^2 (\vartheta/2)} \right),$$

(29)

Hence, (28) gives the representation of a solution to (23) in terms of some vector field $a$ on $S$.

**Theorem 4.3 [14].** For any solution $E$ of the problem (23), (24), the formula (28) with $a = E|_{S_\infty}$ recovers $E$ in the whole exterior domain $\Omega_2$.

Therefore, it follows similarly to the two-dimensional case that the operator in (28) is a projection and the components $\varphi, \tau_2$ and $\tau_3$ can be obtained by means of the values of these functions on $S_\infty$.

The remaining two components $\vartheta_4$ and $\vartheta_1$ are to be found by extending our two-dimensional analysis [1], [3] of the transport equations to three dimensions. If an overlapping domain decomposition is used, then the solution of the transport equations (11), (12) can be represented by

$$\begin{align*}
\vartheta_4(\bar{x}) &= \vartheta(x) + g_1(\bar{x}_2, \bar{x}_3) \\
\vartheta_1(\bar{x}) &= -\vartheta(x) + g_2(\bar{x}_2, \bar{x}_3),
\end{align*}$$

(30)

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where $\bar{p}$ is already known and where, due to (24), $g_1$ and $g_2$ are non-zero only in the “shadow” (wake) domain $\Omega_W$, i.e.

\[
\begin{align*}
    g_1 &\equiv 0 \text{ in } \Omega_2 \setminus \Omega_W, \\
    g_1 &\equiv (\bar{p} - \bar{p})_\omega \text{ in } \Omega_W, \quad \text{and} \\
    g_2 &\equiv 0 \text{ in } \Omega_2 \setminus \Omega_W, \\
    g_2 &\equiv (\bar{p}_1 + \bar{p})_\omega \text{ in } \Omega_W.
\end{align*}
\]

We summarize the results obtained in this section as follows:

**Theorem 4.4.** The general solution to (11), (12) (equivalently (23), (24)) in $\Omega_2$ is expressed by (28) and (30), where $a$ is a function satisfying (26) and where the functions $g_1, g_2$ are defined in (31). If $a$ in (28) is the trace of some solution $E = (\bar{p}, \bar{v}_2, \bar{v}_3)$ to (11), (12) (equivalently (23), (24)) then formula (28) recovers $E$ in $\Omega_2$.

As in the 2D case, the flux condition (20) can be imposed in the opposite direction: $LE \rightarrow NS$. Using this transmission condition and the solution of the linearized Euler equations in the outer domain $\Omega_2$, we obtain far-field boundary conditions at the artificial boundary $S_\infty$. Details of computational aspects are discussed in [14].

References


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